

Here are given results of activation intercomparison in iron and aluminum for short cooling time, (namely $t=1$ d). Different assumptions for irradiation time T were taken -- 30d, 100d, and 10y. For $T=10$ y three cases were calculated – ten irradiation periods with $T_i=120$ d ($t_i=245$ d), $T_i=180$ d ($t_i=185$ d), and $T_i=365$ d ($t_i=0$ d).

Results of simulations in terms of specific Gamma-equivalent are given in Table 1. Total Gamma-equivalent and Gamma-equivalent of dominant radionuclide are presented for the both materials. Methods for estimation and other assumptions are the same as in the previous note of May 15.

Table 1
Specific Gamma-equivalents, $(\text{Sv}\cdot\text{m}^2)/(\text{cm}^3\cdot\text{s})$

Material	Nuclide	T=30 d	T=100 d	T=10y ($T_i=120\text{d}$)	T=10y ($T_i=180\text{d}$)	T=10 y ($T_i=365$ d)
Al	Na-24	2.49E-14	2.49E-14	2.49E-14	2.49E-14	2.49E-14
	tot	2.65E-14	2.99E-14	4.85E-14	5.95E-14	9.05E-14
Fe	Mn-52	5.40E-13	5.60E-13	5.60E-13	5.60E-13	5.60E-13
	tot	7.20E-13	8.10E-13	8.40E-13	8.70E-13	9.20E-13

Table 2
Omega factors, $(\text{Sv}/\text{h})/(\text{Stars}\cdot\text{cm}^{-3}\cdot\text{s}^{-1})$

Material	Author	T=30 d	T=100 d	T=10y ($T_i=120\text{d}$)	T=10y ($T_i=180\text{d}$)	T=10 y
Al	Morev	3.72E-09	4.20E-09	6.81E-09	8.35E-09	1.27E-08
	Huntinen	4.70E-09	-	-	-	1.60E-08
Fe	Morev	1.25E-08	1.41E-08	1.46E-08	1.51E-08	1.60E-08
	Huntinen	4.00E-09	-	-	-	8.20E-09

One can see from the Table 2, that Morev and Huntinen have quite similar results for Al activation. The difference of about 30% may be explained by either different spectra (Mika used real specter, Morev used $E_p=600$ MeV) or cross-sections (Morev use protons cross-section for pions – for very high energies it produce conservative estimation).

Morev's results in Fe are 3 times above for $T=30$ d and 2 times above for $T=10\text{y}$.

So, I conclude that disagreement in Fe/Al ratio for short cooling time results from disagreement in Fe activation, and first of all in Mn^{52} cross-section. In my cross-section library I have only $\text{Mn}^{52\text{m}+\text{g}}$ cross-section and no data on branching ratio for ground and isomer production, consequently I used an ultimately conservative assumption as if the both nuclides are produced independently with branching ratio 1.0 (Mika used 0.5). In addition, the cross-section itself seems to be overestimated by a factor of 2 (the reaction is not very good measured and I used calculated data; all available experimental points lay beneath the calculated curve). An overestimation of Mn^{52} cross-section by factor of 4 may explain the disagreement. Since Mn^{52} half decay is 5.6 days, the difference in omega factors for $t=5-7$ days will be less evident.